La Société Guernesiaise Astronomy Section

Eclipses of the Sun and Moon

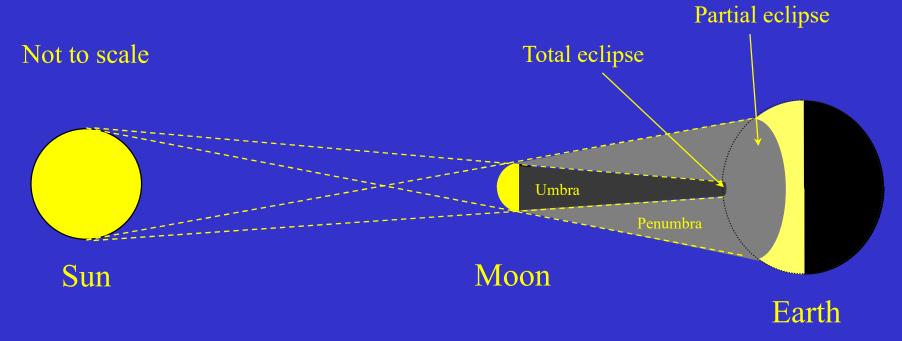
Why they happen, when they happen, and how to observe them safely





The geometry of a solar eclipse

The Moon passes in front of the Sun, blocking it out either totally or partially.



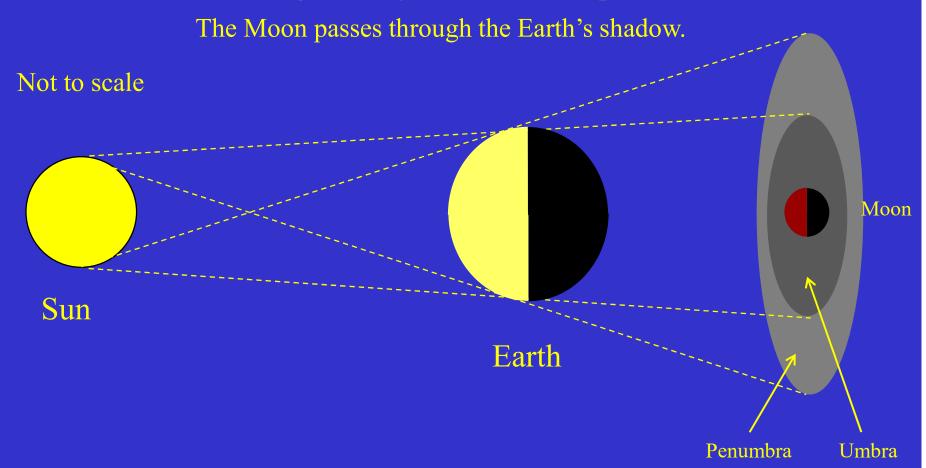
You have to be within the umbra to see the Sun totally eclipsed.

That is within the small spot (typically about 150km, 100 miles wide) on the Earth.

The spot moves rapidly, so totality lasts only a matter of minutes.

For hundreds of miles either side, within the penumbra, there is a partial eclipse.

The geometry of a lunar eclipse

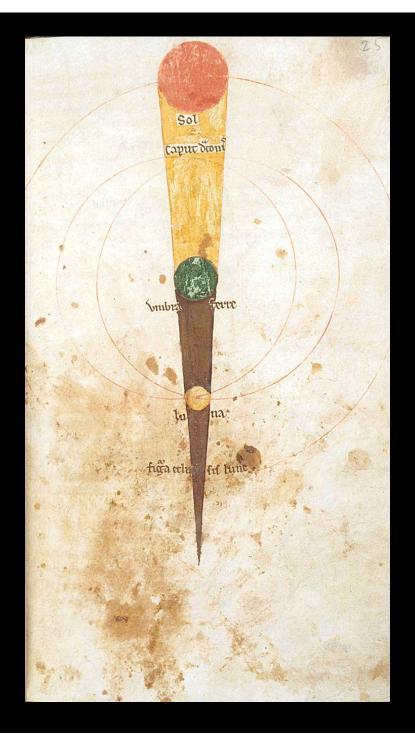


At the Moon's distance the umbra is about 2.5 times the diameter of the Moon. If the Moon passes completely through the umbra it is totally eclipsed.

If it passes partially through the umbra then it is partially eclipsed.

As it passes through the penumbra it is only dimly eclipsed.

Lunar eclipses are visible from anywhere on the night side of the Earth.

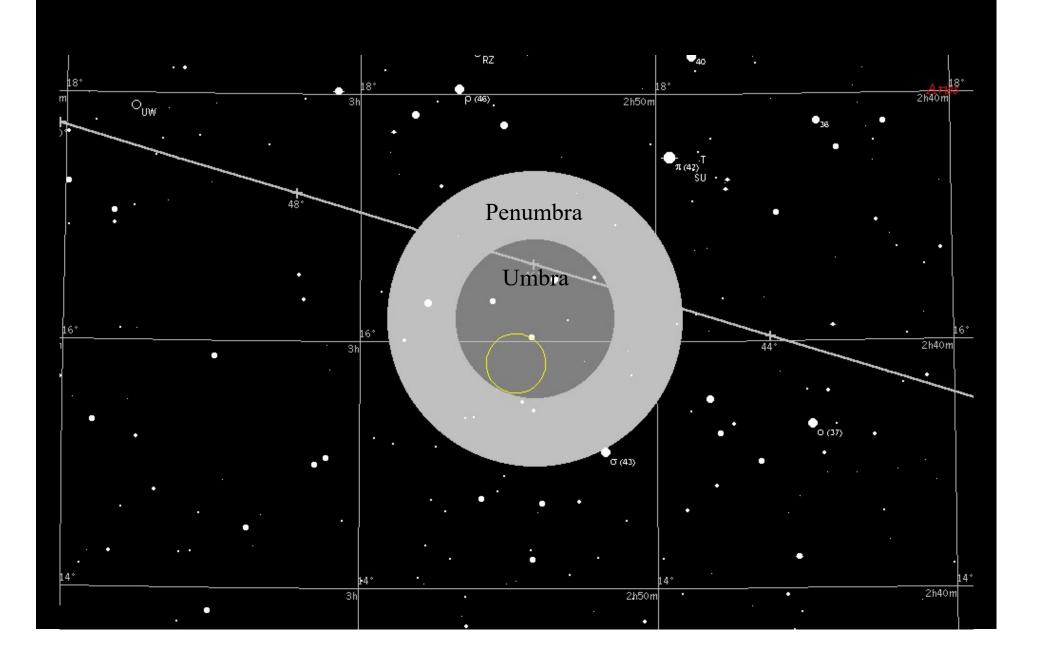


The causes of eclipses were well understood by ancient peoples. The Greek Anaxagoras (c500 BC) reasoned that the Earth was round because of the shape of its shadow on the Moon during a lunar eclipse.

Diagram of a lunar eclipse. Johannes de Sacro Bosco.

England, 13th century.

This graphic shows the Moon (yellow circle) just within the umbra of the Earth's shadow



When the Moon is totally eclipsed it often has a deep red colour.

This is caused by sunlight being refracted (bent) by the Earth's atmosphere and falling onto the Moon.



This montage of five images shows the extent of the Earth's shadow



Here's another sequence of a total lunar eclipse



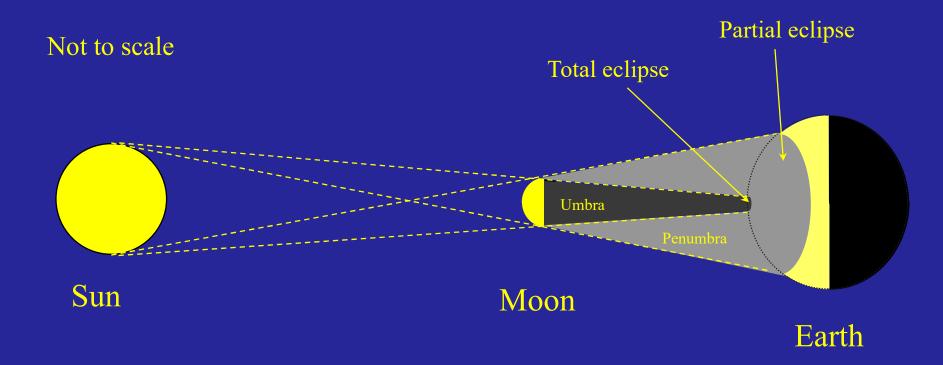


Having foreground objects can produce an impressive picture



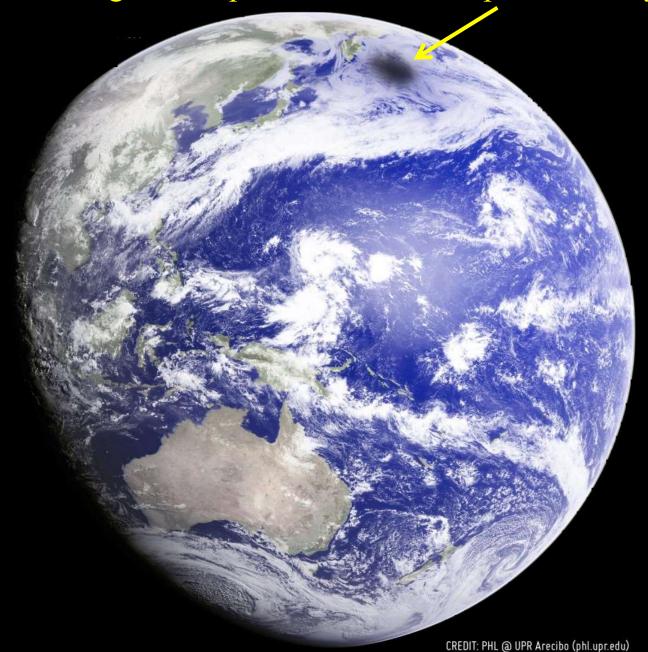


Now let's have a look at solar eclipses



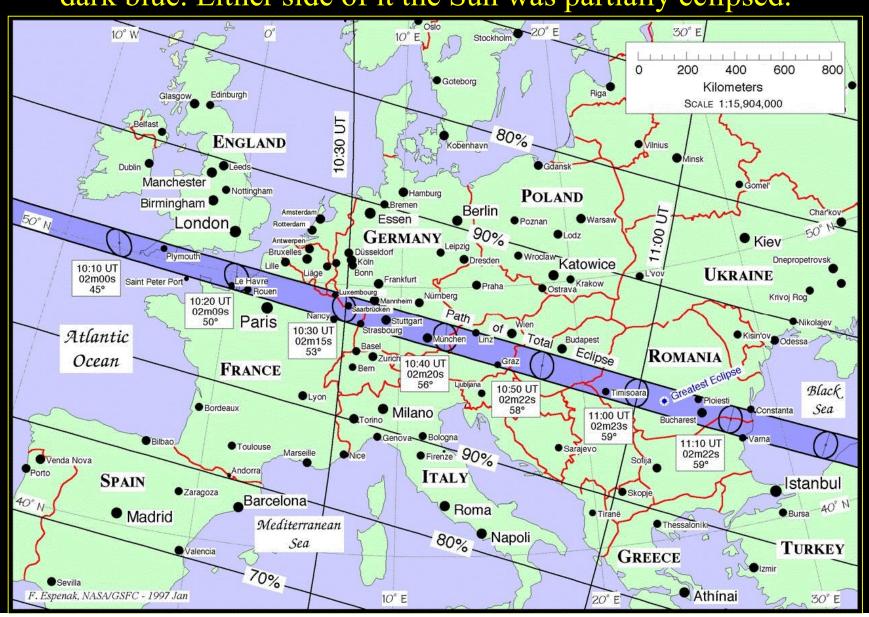
Remember that they happen when the Moon passes directly in front of the Sun, the total eclipse only being visible from a relatively small spot on the Earth.

This image from space shows a dark spot caused by a solar eclipse

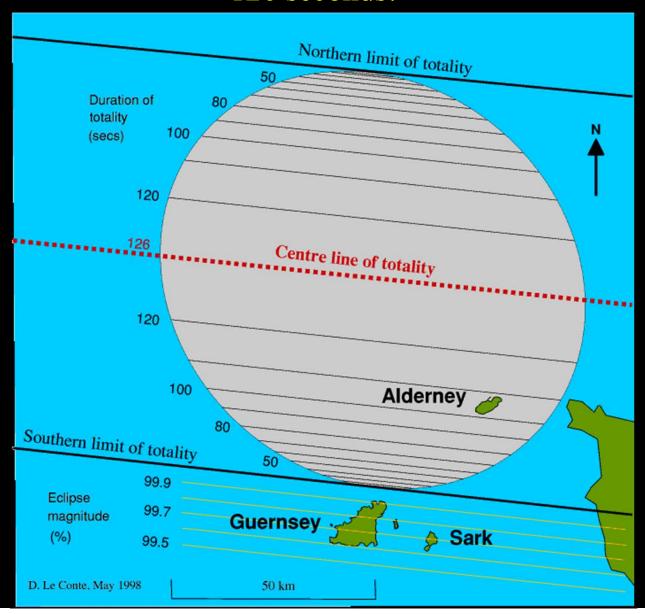


Annular solar eclipse from space (MTSAT in geostationary orbit) 21 May 2012

The spot moves rapidly across the surface of the Earth. Here is the path of the 11 August 1999 eclipse. The total eclipse band is shown in dark blue. Either side of it the Sun was partially eclipsed.



As it passed across the English Channel its speed was over 3000 kph (1900 mph) from west to east. At the centre line totality lasted just 126 seconds!



WARNING

It is dangerous to look directly at the Sun, especially with any optical instrument such as binoculars or a telescope. Blindness can result.

Only specialised equipment or projection methods should be used.

One of the easiest ways for <u>non-telescopic</u> viewing is to use special approved eclipse glasses.

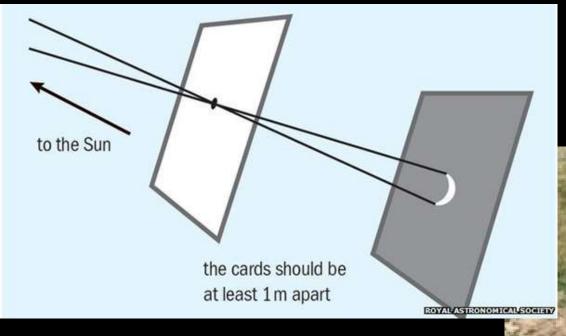
Put the glasses on before looking at the Sun.



Welders goggles with filters rated 14 or higher are also safe.



Another method is pinhole projection. Make a small hole in a piece of card and hold it with your back to the Sun so that an image of the Sun is projected onto another piece of white card about a metre away. Do not look through the hole!



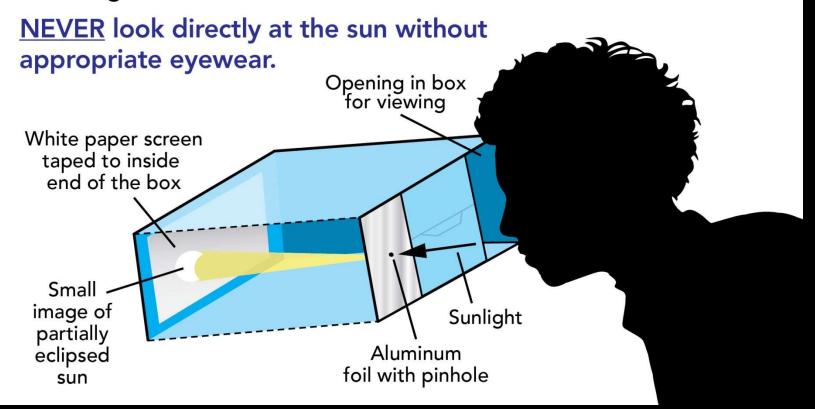


You can make the image easier to see by using a box, such as a cereal box to make a simple eclipse projector

MAKE YOUR OWN ECLIPSE PROJECTOR

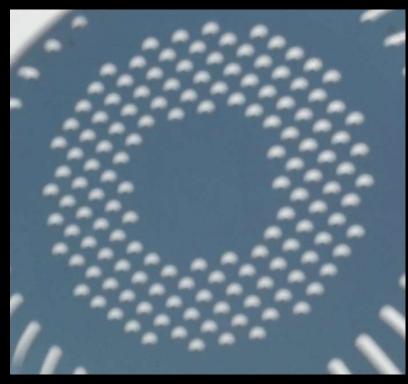
You can make this simple eclipse projector with almost any cardboard box, paper, tape and foil.

The longer the distance from the pinhole to screen, the larger the image of the sun will be.



Anything with a hole in it can be used, even a hat! Here a kitchen colander produces multiple solar images during an eclipse.

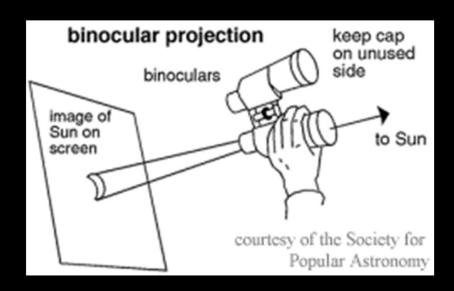




You don't even need any equipment at all if there are trees around. The gaps between leaves form pinhole images of the Sun on the ground, and these become crescent-shaped during a solar eclipse.



Larger, detailed images can be obtained by projecting sunlight through binoculars or a refracting telescope.



In all these projection methods you stand with your <u>back</u> to the Sun.



No filters that you wear should be used for looking through telescopes, binoculars or cameras.

For these, special optical filters which fit over the front of the instrument's main lens are needed.

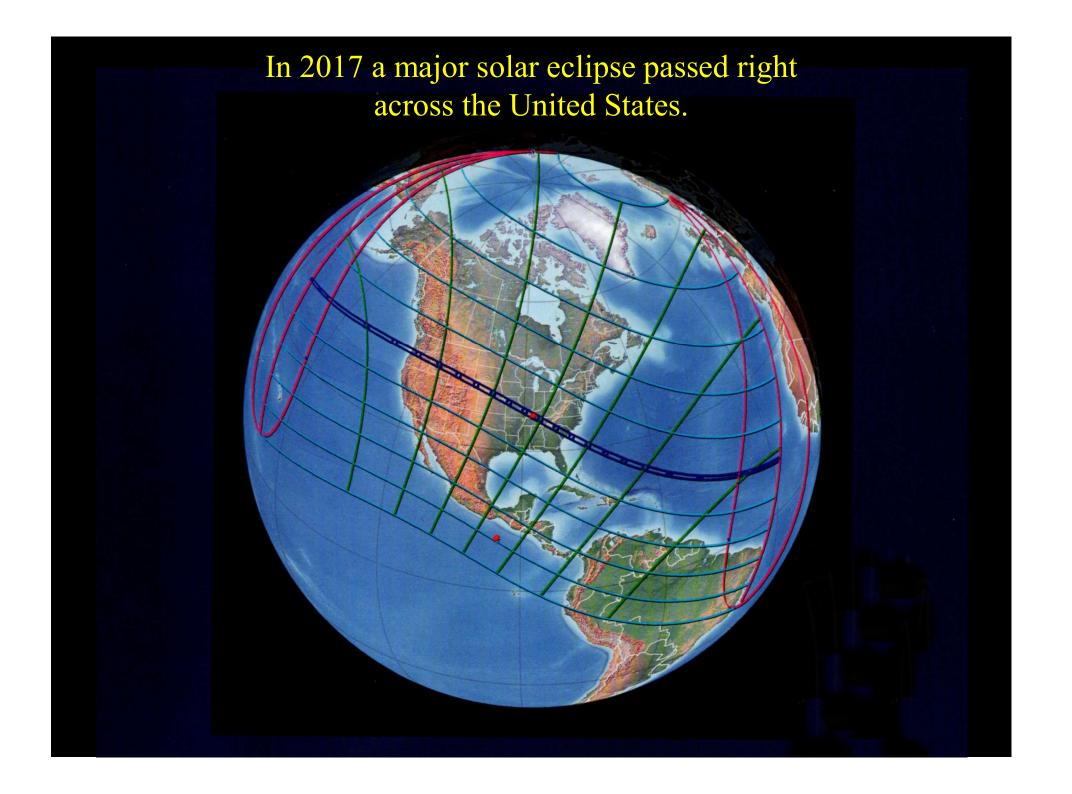


All these methods are fully described at www.mreclipse.com/Totality3/TotalityCh11.html

Here is a partial solar eclipse taken with an 11-inch reflecting telescope with a proper filter.



David Le Conte, Guernsey 12 October 1996



The appearance of the corona is affected by the Sun's magnetic field



David Le Conte / Jean Dean

Click to continue



This montage shows the entire sequence of the solar eclipse



As totality approaches things start getting really exciting. The sky gets noticeably darker, shadows sharpen, it will feel cooler, animals may start acting strangely, there may possibly be a wind, and bright planets and stars may become visible.

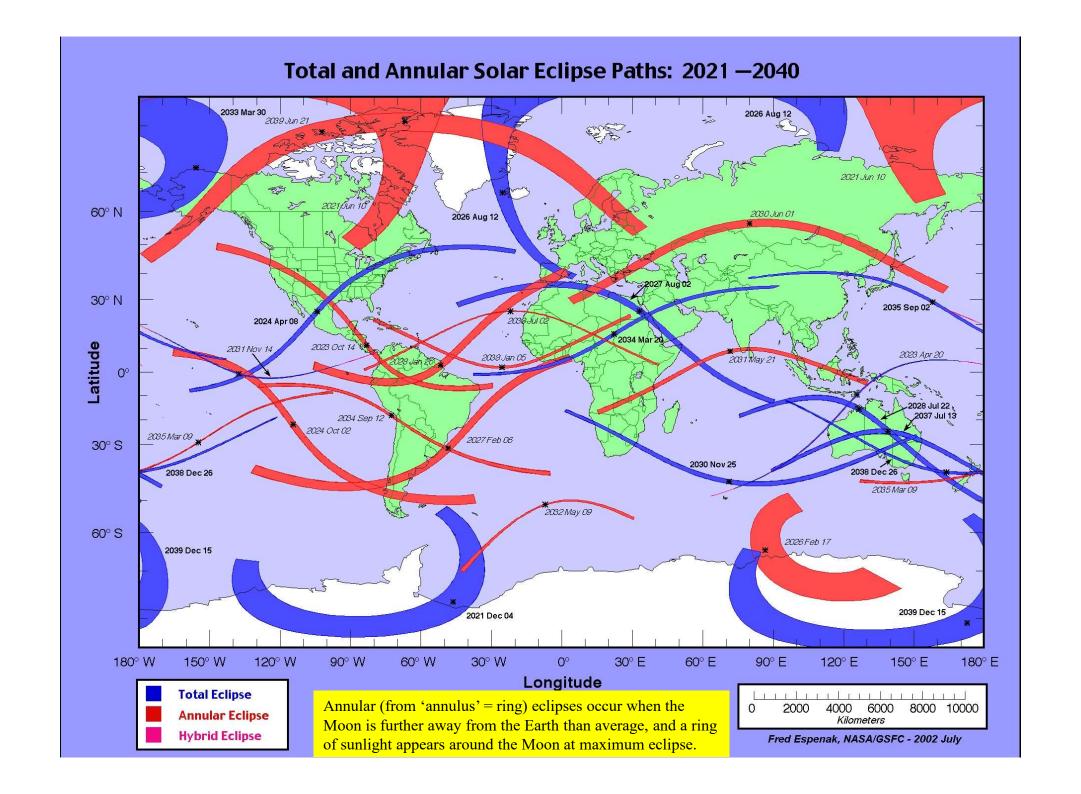
In the last couple of minutes before totality look out for 'shadow bands' – rippling bands seen on the light ground surface or a white sheet. They are caused by light from the thin crescent Sun affected by turbulence in the Earth's atmosphere.

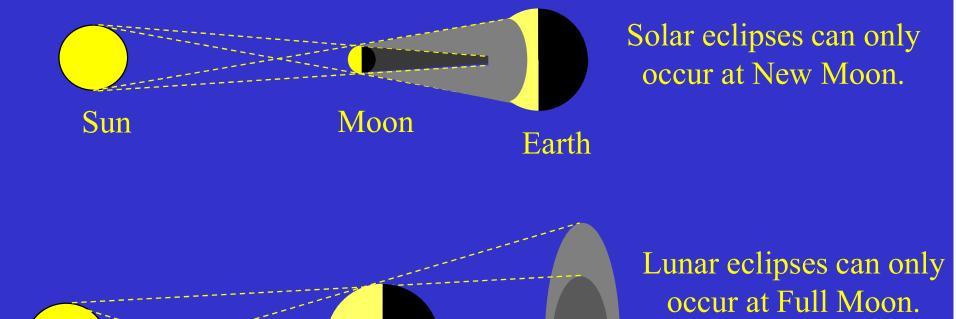
In the 30 seconds or so before totality look towards the west. You may see the Moon's shadow racing towards you!

If you have eclipse glasses you <u>must</u> use them during the partial phases of the eclipse. But in order to see the solar corona you must remove them during the few minutes of totality, using them again as soon as the Sun starts reappearing.

During totality note that all of the horizon appears to have the colours of sunset $-a 360^{\circ}$ twilight!

The longest totality can last is 7½ minutes, but most are only 2-4 minutes!





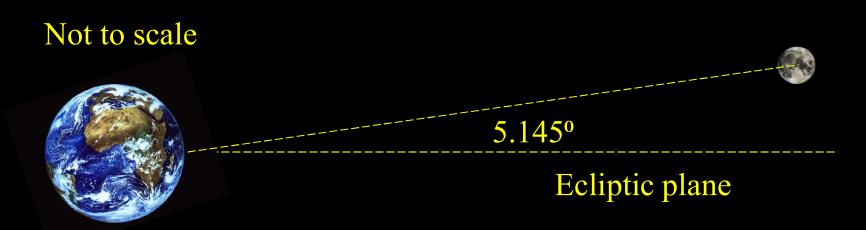
Moon

So why aren't there eclipses at every New and Full Moon?

Earth

Sun

The answer is that the Moon's orbit is inclined at just over 5° to the plane of the Earth's orbit around the Sun (called the 'Ecliptic').



So the Moon usually passes above or below the Sun and the Earth's shadow. Only rarely do they line up to cause an eclipse.

The pattern of eclipses repeats every 18 years, 11 days, 8 hours (a period called the Saros).

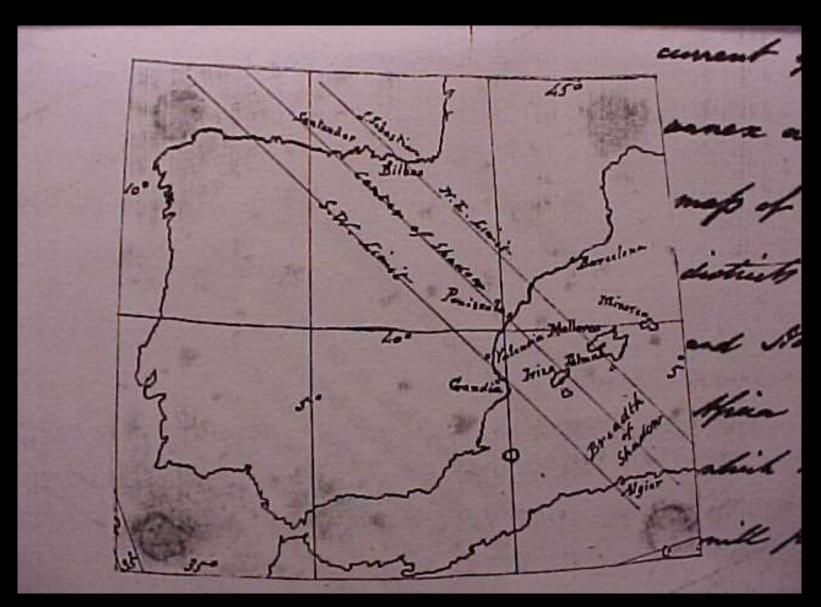
There are between 2 and 5 solar eclipses (not necessarily total) every year, and between 0 and 3 lunar eclipses (ignoring penumbral eclipses).

Some Guernsey connections with historic solar eclipses.

In 1858 Guernseyman Thomas Lukis Mansell took the first known astronomical photographs in Guernsey – of a solar eclipse.

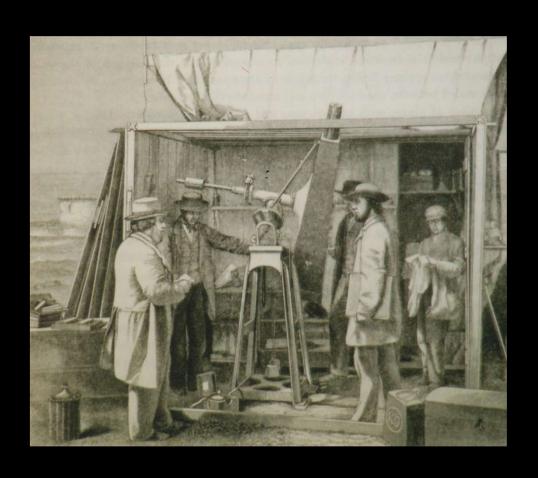


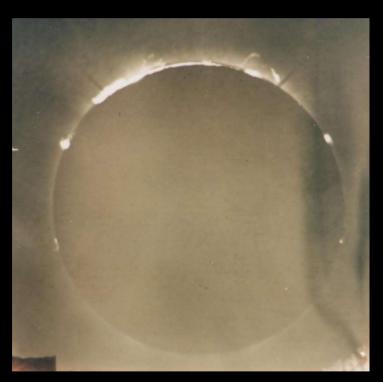
In 1860 there was a total solar eclipse in the north of Spain. Here is the Astronomer Royal's plan of the eclipse path.



Guernseyman Warren De La Rue was on the Astronomer Royal's eclipse expedition ...

... and took the first scientific pictures of an eclipse, which showed that the prominences were part of the Sun.







In 1870 Guernsey artist Paul Jacob Naftel was invited to join the Royal Astronomical Society's eclipse expedition to southern Spain ...

... and painted the scene to show the details of the corona and the colours of the landscape.



This presentation has been brought to you by the Astronomy Section of La Société Guernesiaise, Guernsey, in the British Channel Islands.

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